

# Parallel Space-Time Likelihood Optimization for Air Pollution Prediction on Large-Scale Systems

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# I. Motivation

Modern-Era Retrospective Analysis for Research and Applications, version 2 (MERRA-2) reanalysis  
log particulate matter (PM) data on January 1, 2016

Fig. 1: Middle East

Fig. 2: United States



# Overview

- I Motivation
- II Preliminaries
- III Contributions
- IV Two-Level Parallelization Framework
- V Performance Results & Analysis
- VI Summary



## II. Preliminaries



## II. Preliminaries: Space-Time Geostatistics

- ▶ Suppose  $\mathbf{Z}_1 \in \mathbb{R}^{n_1}$  is a vector of measurements from  $n_1$  sampled space-time locations, i.e.,  
 $\mathbf{Z}_1 = \{Z(\mathbf{s}_1, t_1), \dots, Z(\mathbf{s}_{n_1}, t_{n_1})\}^\top$ , where  $Z(\mathbf{s}, t) \in \mathbb{R}$  is the measurement at sampled space-time location  $(\mathbf{s}, t) \in \mathbb{R}^d \times \mathbb{R}$ .
- ▶ Suppose  $\mathbf{Z}_2 \in \mathbb{R}^{n_2}$  is a vector of missing measurements at  $n_2$  unsampled space-time locations, i.e.,  
 $\mathbf{Z}_2 = \{Z(\tilde{\mathbf{s}}_1, \tilde{t}_1), \dots, Z(\tilde{\mathbf{s}}_{n_2}, \tilde{t}_{n_2})\}^\top$ , where  $Z(\tilde{\mathbf{s}}, \tilde{t}) \in \mathbb{R}$  is the missing measurement at unsampled space-time location  $(\tilde{\mathbf{s}}, \tilde{t}) \in \mathbb{R}^d \times \mathbb{R}$ .
- ▶ Suppose  $\mathbf{Z}_1$  and  $\mathbf{Z}_2$  are **jointly Gaussian**, i.e.,

$$\begin{bmatrix} \mathbf{Z}_1 \\ \mathbf{Z}_2 \end{bmatrix} \sim \mathcal{N}_{n_1+n_2} \left( \begin{bmatrix} \boldsymbol{\mu}_1 \\ \boldsymbol{\mu}_2 \end{bmatrix}, \begin{bmatrix} \boldsymbol{\Sigma}_{11} & \boldsymbol{\Sigma}_{12} \\ \boldsymbol{\Sigma}_{21} & \boldsymbol{\Sigma}_{22} \end{bmatrix} \right), \quad (1)$$

where  $\boldsymbol{\mu}_1 \in \mathbb{R}^{n_1}$  and  $\boldsymbol{\mu}_2 \in \mathbb{R}^{n_2}$  are the mean vectors of  $\mathbf{Z}_1$  and  $\mathbf{Z}_2$ , respectively,  $\boldsymbol{\Sigma}_{11}$  and  $\boldsymbol{\Sigma}_{22}$  are the covariance matrices of  $\mathbf{Z}_1$  and  $\mathbf{Z}_2$ , respectively, and  $\boldsymbol{\Sigma}_{12} = \boldsymbol{\Sigma}_{21}^\top$  is the cross-covariance matrix of  $\mathbf{Z}_1$  and  $\mathbf{Z}_2$ .



## II. Preliminaries: Space-Time Geostatistics

- Prediction for  $\mathbf{Z}_2$  can be performed using the concept of **conditional distribution of Gaussian** processes, i.e.,

$$\mathbf{Z}_2 | \mathbf{Z}_1 \sim \mathcal{N}_{n_2} \{ \boldsymbol{\mu}_2 + \boldsymbol{\Sigma}_{21} \boldsymbol{\Sigma}_{11}^{-1} (\mathbf{Z}_1 - \boldsymbol{\mu}_1), \boldsymbol{\Sigma}_{22} - \boldsymbol{\Sigma}_{21} \boldsymbol{\Sigma}_{11}^{-1} \boldsymbol{\Sigma}_{12} \}. \quad (2)$$

This means that the best prediction for the missing measurements vector  $\mathbf{Z}_2$ , denoted  $\hat{\mathbf{Z}}_2$ , is

$$\hat{\mathbf{Z}}_2 = \boldsymbol{\mu}_2 + \boldsymbol{\Sigma}_{21} \boldsymbol{\Sigma}_{11}^{-1} (\mathbf{Z}_1 - \boldsymbol{\mu}_1). \quad (3)$$

- Prediction relies on the mean vectors and covariance matrices.
- **In practice**, mean vectors and covariance matrices are estimated from the given measurements,  $\mathbf{Z}_1$ .
- This is done by choosing and fitting a **parametric** mean function and a **parametric** covariance function to the data.



## II. Preliminaries: Space-Time Geostatistics

- ▶ Suppose  $\mu_1 = \mu_2 = \mathbf{0}$ .
- ▶ To model the covariance matrices  $\Sigma_{11}$ ,  $\Sigma_{22}$ , and  $\Sigma_{12}$ , we need a parametric space-time covariance function, denoted  $C(\mathbf{s}_1, \mathbf{s}_2; t_1, t_2 | \theta)$ , such that

$$\Sigma_{11}(\theta) = \{C(\mathbf{s}_i, \mathbf{s}_j; t_i, t_j | \theta)\}_{i,j=1}^{n_1},$$

$$\Sigma_{22}(\theta) = \{C(\tilde{\mathbf{s}}_i, \tilde{\mathbf{s}}_j; \tilde{t}_i, \tilde{t}_j | \theta)\}_{i,j=1}^{n_2}, \text{ and}$$

$$\Sigma_{12}(\theta) = [\{C(\mathbf{s}_i, \tilde{\mathbf{s}}_j; t_i, \tilde{t}_j | \theta)\}_{i=1}^{n_1}]_{j=1}^{n_2}.$$

- ▶  $C(\mathbf{s}_1, \mathbf{s}_2; t_1, t_2 | \theta)$  describes the **strength of dependence** of the measurements at any two space-time locations  $(\mathbf{s}_1, t_1)$  and  $(\mathbf{s}_2, t_2)$ .



## II. Preliminaries: Space-Time Geostatistics

- ▶ A popular space-time covariance function proposed in **Gneiting (2002)** has the form:

$$C(\mathbf{h}, u | \boldsymbol{\theta}) = \frac{\sigma^2}{|u|^{2\alpha}/a_t + 1} \mathcal{M}_\nu \left\{ \frac{\|\mathbf{h}\|/a_s}{(|u|^{2\alpha}/a_t + 1)^{\beta/2}} \right\},$$

where  $\mathbf{h} = \mathbf{s}_1 - \mathbf{s}_2$  and  $u = t_1 - t_2$ ,  $\mathcal{M}_\nu$  is the univariate Matérn correlation function with parameter vector  $\boldsymbol{\theta} = (\sigma^2, a_s, \nu, a_t, \alpha, \beta)^\top \in \mathbb{R}^6$ , such that  $\sigma^2 > 0$  is the variance parameter,  $\nu > 0$  and  $\alpha \in (0, 1]$  are the smoothness parameters in space and time, respectively,  $a_s, a_t > 0$  are the range parameters in space and time, respectively, and  $\beta \in [0, 1]$  is the space-time interaction parameter.

- ▶ When  $\beta = 0$ , **separable** model
- ▶ When  $\beta > 0$ , **nonseparable** model





## II. Preliminaries: Space-Time Geostatistics

Simulated space-time realizations  
from the space-time covariance function model

Fig. 3: Nonseparable

Fig. 4: Separable



## II. Preliminaries: Spatio-Temporal Geostatistics

### Gaussian log-likelihood function

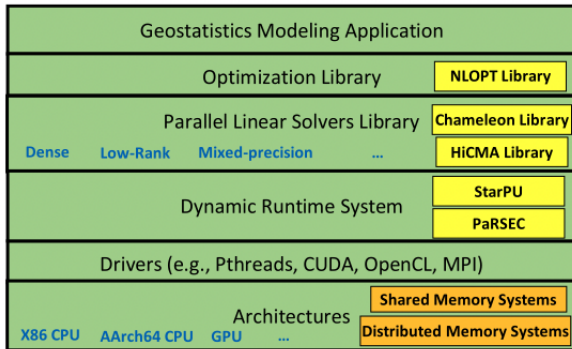
$$l(\boldsymbol{\theta}) = -\frac{n}{2} \log(2\pi) - \frac{1}{2} \log |\boldsymbol{\Sigma}(\boldsymbol{\theta})| - \frac{1}{2} \mathbf{Z}^\top \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \mathbf{Z},$$

Here  $|\boldsymbol{\Sigma}(\boldsymbol{\theta})|$  is the determinant of  $\boldsymbol{\Sigma}(\boldsymbol{\theta})$ .

The grand challenge in large-scale Gaussian geostatistical modeling lies in the inversion of  $\boldsymbol{\Sigma}(\boldsymbol{\theta})$ .



## II. Preliminaries: The ExaGeoStat Software



**Fig. 5:** The ExaGeoStat software layers for geostatistics applications.  
Source: Huang et al. (2021)



## II. Preliminaries: State-of-the-Art Dense Linear Algebra Libraries<sup>12/33</sup>

### Libraries

Data layout formats

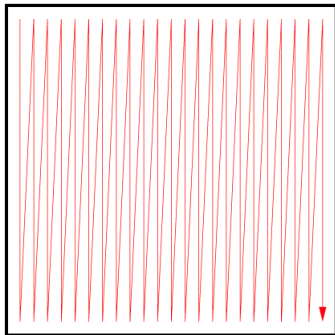


Fig. 6: LAPACK: Column-major data layout format.

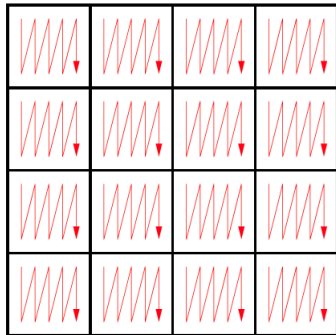


Fig. 7: Chameleon: Tile data layout format.



## II. Preliminaries: Directed Acyclic Graph (DAG)

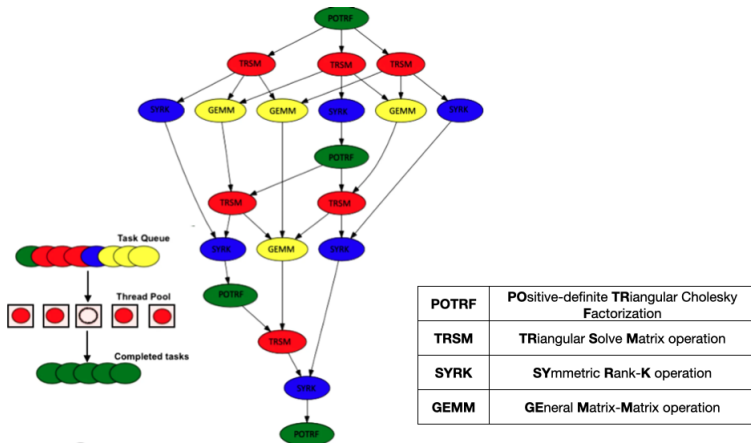


Fig. 8: An example of task-based DAG to perform Cholesky factorization of 4-tiles by 4-tiles matrix.



## II. Preliminaries: Particle Swarm Optimization

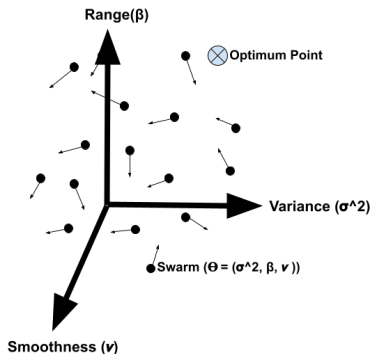


Fig. 9: Iteration 0

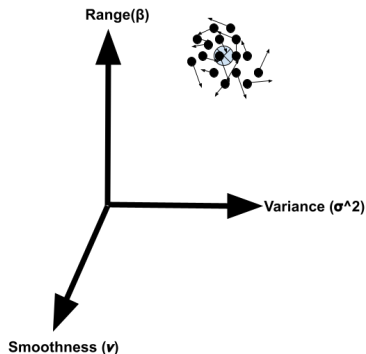


Fig. 10: Iteration  $N$

PSO-based MLE optimization of a univariate purely spatial covariance function.



### III. Contributions



### III. Contributions

1. **Implementation** of high-performance space-time model on large-scale systems
2. **Visualization** of space-time random fields generated by the high-performance implementation
3. **Incorporation** of the PSO algorithm to the MLE operation to utilize the execution performance on distributed environments
4. **Illustration** of benefits of flexible vs. simple space-time model via large scale space-time experiments
5. **Application** to air pollution datasets from the Middle East and US





## IV. The Proposed Two-Level Parallelization Framework



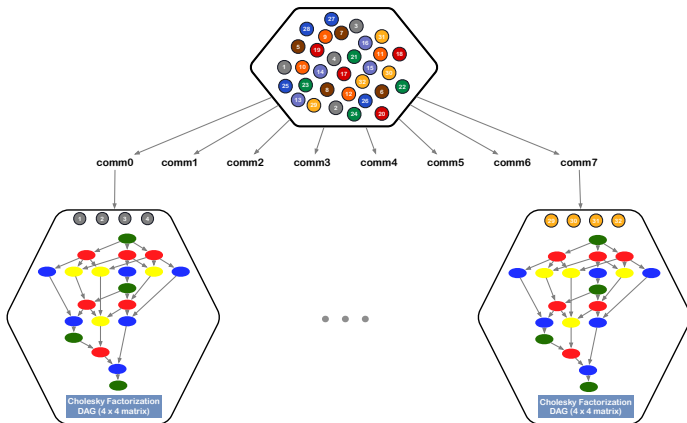
## IV. Proposed: Parallel Optimization Strategy

- ▶ ***MPI\_COMM\_WORLD***: MPI default communicator
- ▶ ***MPI\_Comm\_split***: partition the default communicator into disjoint subgroups associated with different sub-communicators
- ▶ PPSwarm algorithm

We split the default communicator into a set of sub-communicators where each can be used to evaluate a single log-likelihood function solution.



## IV. Proposed: Framework



**Fig. 11:** Testcase using 32 nodes and 8 MPI sub-communicators. Each sub-communicator includes 4 nodes that estimate the log-likelihood function with a certain set of parameters in parallel using the StarPU runtime system.



## V. Performance Results & Analysis

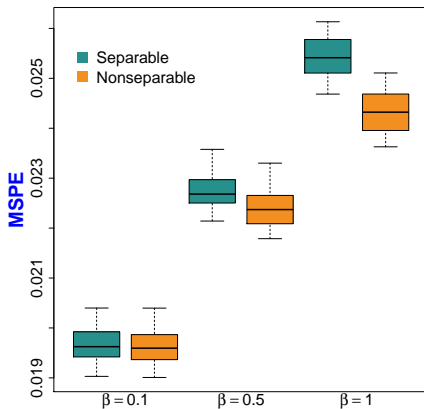


## V. Performance Results & Analysis

- ▶ **Synthetic data** from the space-time model with  $\beta \in \{0.1, 0.5, 1\}$ 
  - ▶ 400 spatial locations
  - ▶ 100 temporal locations
  - ▶  $40,000 \times 40,000$  covariance matrix size
- ▶ **Real data** of particulate matter
  - ▶ 550 spatial locations
  - ▶ 730 temporal locations
  - ▶  $401,500 \times 401,500$  covariance matrix size
- ▶ The performance is tested on an Intel-based Cray XC40 system with 6,174 compute nodes, each of which has two 16-core Intel Haswell CPUs at 2.30 GHz and 128 GB of memory.
- ▶ All the experiments were conducted on the whole number of cores with different nodes.



## V. Performance Results & Analysis: Synthetic Data



**Fig. 12:** Boxplots of the prediction errors when fitting a separable and nonseparable model on space-time data with varying degrees of space-time interactions. Weak, moderate, and strong space-time interactions are represented by  $\beta = 0.1, 0.5$ , and  $1$ , respectively.



# V. Performance Results & Analysis: Synthetic Data

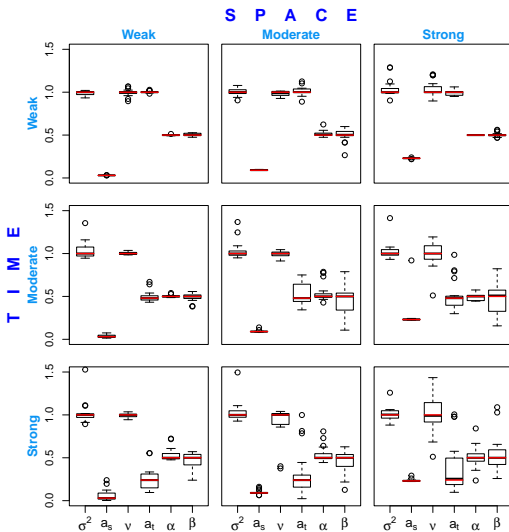
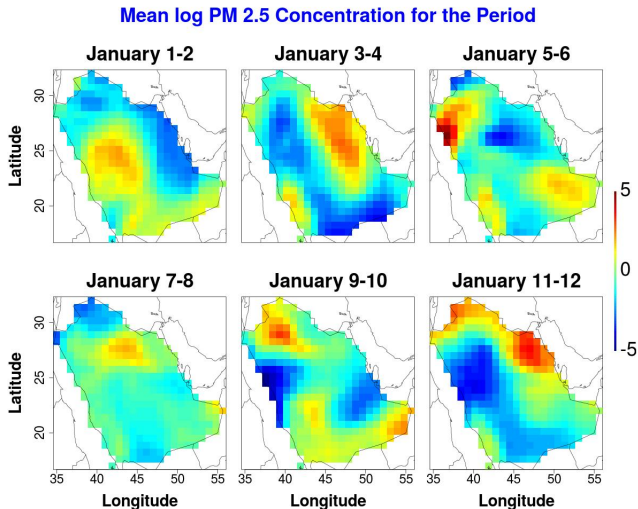


Fig. 13: Boxplots of parameter estimates under varying degrees of space-time dependence. The true parameters are highlighted in red.

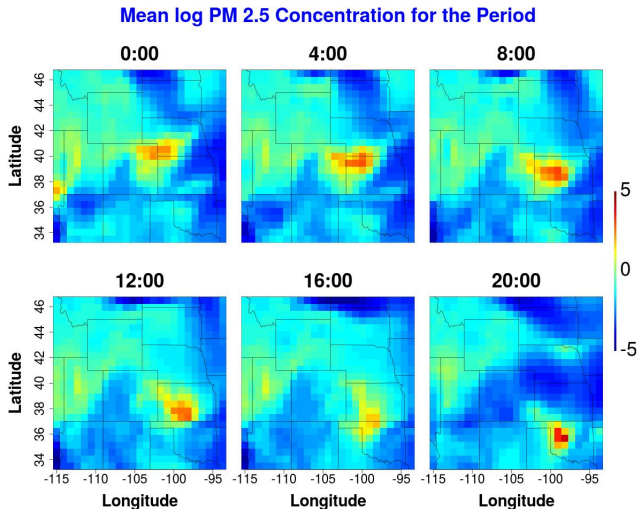
## V. Performance Results & Analysis: Real Data



**Fig. 14:** Visualization of the log PM<sub>2.5</sub> dataset after space-time mean removal at the first six time points in 2016 over Saudi Arabia.



## V. Performance Results & Analysis: Real Data



**Fig. 15:** Visualization of the log PM<sub>2.5</sub> dataset after space-time mean removal at four hour intervals on January 1, 2016 over the Midwest US.

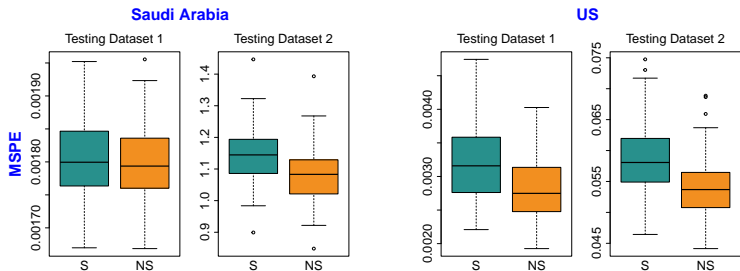
## V. Performance Results & Analysis: Real Data

**Table 1:** A summary of the estimated parameters of the nonseparable (NS) and separable (S) models and their corresponding errors (MSPE) and prediction uncertainty (PU) for the Saudi Arabia and US datasets. MSPE<sub>1</sub> and PU<sub>1</sub> correspond to Testing Dataset 1, while MSPE<sub>2</sub> and PU<sub>2</sub> point to Testing Dataset 2. The best model reports the lower MSPE and PU.

	Model	$\hat{\sigma}^2$	$\hat{a}_s$	$\hat{v}$	$\hat{a}_t$	$\hat{\alpha}$	$\hat{\beta}$	MSPE <sub>1</sub> / PU <sub>1</sub>	MSPE <sub>2</sub> / PU <sub>2</sub>
Saudi Arabia	NS	1.29	1.34	2.15	1.12	0.14	0.75	0.0017/76	1.08/875
	S	2.61	1.27	2.15	2.04	0.03	0	0.0018/78	1.14/1066
US	NS	0.47	1.33	1.12	6.77	0.72	0.14	0.0028/155	0.05/322
	S	2.12	1.54	1.47	7.99	0.48	0	0.0031/134	0.06/1118



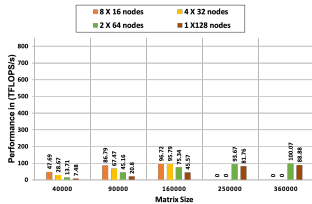
## V. Performance Results & Analysis: Real Data



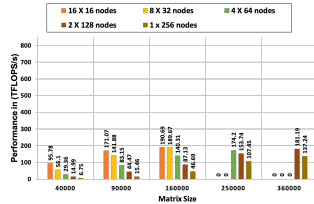
**Fig. 16:** Boxplots of the prediction errors for the separable (S) and nonseparable (NS) models in the real data pseudo cross-validation study.



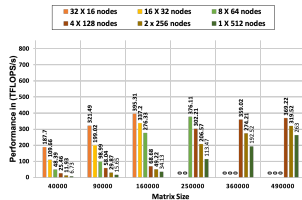
# V. Performance Results & Analysis: Real Data



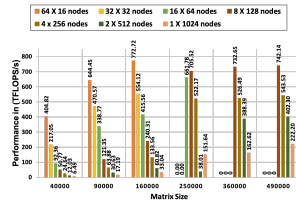
(a) 128 nodes.



(b) 256 nodes.



(c) 512 nodes.

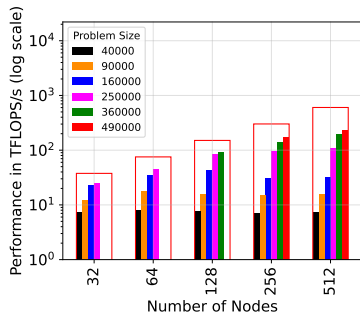


(d) 1024 nodes.

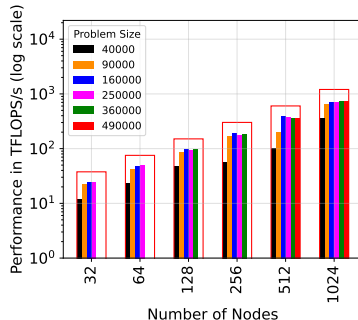
Fig. 17: Performance of a single MLE optimization step using different number of nodes on Shaheen-II Cray XC40 Supercomputer.



## V. Performance Results & Analysis: Real Data



(a) One-level parallelization.



(b) Two-level parallelization.

**Fig. 18:** Performance of one MLE optimization step using single and  $n$  MPI communicators on Shaheen-II Cray XC40 Supercomputers. In (b)  $x$  MPI sub-communicators is used where  $x$  is tuned for performance.



# V. Performance Results & Analysis: Real Data

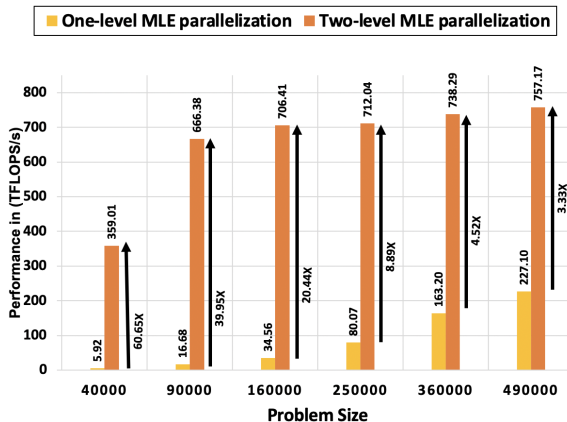
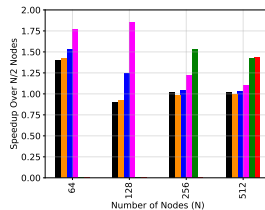
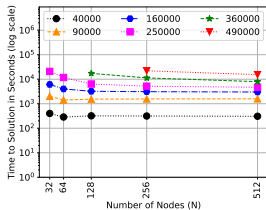


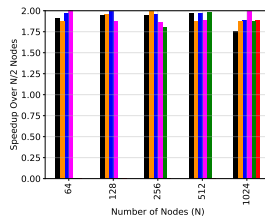
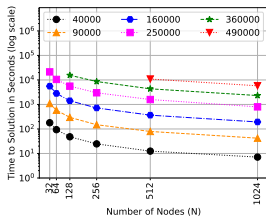
Fig. 19: One-level versus two-level MLE parallelization performance using 1024 nodes on Shaheen-II Cray XC40 system.



# V. Performance Results & Analysis: Real Data



(a) One-level parallelization.



(b) Two-level parallelization.

**Fig. 20:** Time-to-solution of full MLE operation using 100 optimization iterations on Shaheen-II Cray XC40 system. In (b), we tune  $x$  MPI sub-communicators for performance.



## VI. Summary





## VI. Summary

- ▶ Proposed a two-level parallelization framework of geostatistical space-time modeling
  - ▶ Upper level: MPI sub-communicators perform independent log-likelihood function evaluation with different sets of parameters via PPSO algorithm
  - ▶ Inner level: task-based parallel technique is used to perform linear solver operations on a given set of nodes representing a single MPI sub-communicator
- ▶ Demonstrated, through synthetic and real datasets, the merits of the proposed implementation
- ▶ Achieved high prediction accuracy with up to 757 TFLOPS/s using 1024 nodes on the KAUST Shaheen-II Cray XC40 system (around 63% of the theoretical peak)
- ▶ Other avenues for research: **tile low rank** and **mixed-precision approximations** to accelerate further the modeling process



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## References II

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Questions?